



Effective elastic thickness of South America and its implications for intracontinental deformation

3 M. Pérez-Gussinyé

4 Institute of Earth Sciences "Jaume Almera," CSIC, Lluis Sole i Sabaris s/n, Barcelona E-08028, Spain 5 (mperez@ija.csic.es)

6 A. R. Lowry

7 Department of Geology, Utah State University, 4505 Old Main Hill, Logan, Utah 84322, USA

8 A. B. Watts

9 Department of Earth Sciences, University of Oxford, Oxford OX1 3PR, UK

[1] The flexural rigidity or effective elastic thickness of the lithosphere, T_e , primarily depends on its 10 thermal gradient and composition. Consequently, maps of the lateral variability of T_e in continents reflect 11 their lithospheric structure. We present here a new T_e map of South America generated using a compilation 12of satellite-derived (GRACE and CHAMP missions) and terrestrial gravity data (including EGM96 and 13 SAGP), and a multitaper Bouguer coherence technique. Our T_e maps correlate remarkably well with other 14proxies for lithospheric structure: areas with high T_e have, in general, high lithospheric mantle shear wave 15velocity and low heat flow and vice versa. In this paper we focus on the T_e of the stable platform. We find 16 that old cratonic nuclei (mainly Archean and Early/Middle Proterozoic) have, in general, high T_e (>70 km), 17 while the younger Patagonian Phanerozoic terrane has much lower T_e (20–30 km), suggesting that T_e is 18 related to terrane age as has already been noted in Europe. Within cratonic South America, T_e variations are 19observed at regional scale: relatively lower T_e occurs at sites that have been repeatedly reactivated 20throughout geological history as major sutures, rift zones, and sites of hot spot magmatism. Today, these 21 low T_e areas are surrounded by large cratonic nuclei. They concentrate most of the intracontinental 22 seismicity and exhibit relatively high surface heat flow and low seismic velocity at 100 km depth. This 23 implies that intracontinental deformation focuses within relatively thin, hot, and hence weak lithosphere, 24 that cratonic interiors are strong enough to inhibit tectonism, and that the differences in lithospheric 25rigidity, structure, and composition between stable cratons and sites of intracontinental deformation are not 26transient, and may have been maintained, in some cases, for at least 500 m.y. 27

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37 1. Introduction

[2] Continents form by the amalgamation of 38 39 terranes that have stabilized at different times during Earth's history. In general, old terranes 40 (>1.8 Ga) have a lithosphere that is thicker, more 41depleted in basaltic constituents and, consequently, 42more dehydrated than younger ones [Jordan, 1978; 43O'Reilly et al., 2001; Hirth and Kohlstedt, 1996]. 44 The combination of these factors is thought to 45make continental cratonic interiors more resistant 46to subsequent deformation [e.g., Pollack, 1986; 47 Hirth and Kohlstedt, 1996]. 48

[3] A measure of the resistance to vertical defor-49mation of the lithosphere is its flexural rigidity or, 50equivalently, its effective elastic thickness, T_e . A 51recent map of T_e in Europe, for example, shows 52that in this continent areas of high T_e are located 53within older terranes (>1.5 Gyr) where the litho-5455spheric thickness, as inferred from shear wavevelocities and thermal gradient, is much greater 56than in the younger terranes [Pérez-Gussinyé and 57*Watts*, 2005]. Since large T_e correlates well with 58areas where the seismic and thermal lithosphere is 59thick and vice versa, T_e maps of continents may be 60 used not only to better understand their mechanical 61 properties, but also to image the lateral variability 62in their structure. 63

[4] Earth structure is commonly imaged using 64 seismic velocities as a proxy for rock temperature 65and composition. On the other hand, sensitivity 66 analysis of T_e , indicates that it primarily depends 67 on crustal thickness (a compositional feature) and 68 parameters of power law creep, i.e., temperature, 69 composition and to a lesser degree strain rate 70 [Lowry and Smith, 1995; Burov and Diament, 71 1995; Lowry et al., 2000; Brown and Phillips, 722000]. Thus T_e mainly depends on temperature 73 and composition and could be used, in principal, 74 to map lithospheric structure in an analogous way 75as seismic velocities are used. However, while 76 seismic tomography can provide an image of the 77 subsurface at different depths, T_e represents a depth 78integral of physical properties over the lithosphere. 79Despite this, T_e affords a view of the lithosphere 80 that complements tomography because it more 81 directly reflects the rheological strength than does 82 seismic velocity. Moreover, mapping lithospheric 83 structure at continent-wide scale using flexural 84 rigidity has become much simpler because of the 85 recent availability of relatively high resolution 86 (1 degree) global grids of satellite-derived free-air 87

gravity data (from GRACE and CHAMP) that can 88 be combined with terrestrial data. 89

[5] Traditionally, T_e is estimated using either 90 forward (space-domain based) or inverse (spectral- 91 domain based) methods, both of which use topogra- 92 phy and gravity anomalies as input data [*Watts*, 2001]. 93 Exact comparison of the results obtained from differ- 94 ent regional T_e studies is not necessarily straightfor- 95 ward, however, as T_e estimates are sensitive to the 96 dimensions of the data window, the power spectral 97 estimator and the loading model used. 98

[6] Loading models include those that assume only 99 surface loads such as the thrust sheets in orogenic 100 belts, and those that consider both surface and 101 subsurface loads, as for example intracrustal thrusts 102 and magmatic underplating. Models that only 103 include surface loads were used during the late 104 70s and early 80s. While these studies yielded 105 reasonable results for oceanic lithosphere [Watts 106 et al., 1980], they produced very low estimates for 107 cratonic continental lithosphere (e.g., 5–10 km for 108 the cratonic United States [Banks et al., 1977]). 109 Forsyth [1985], however, noted that if subsurface 110 loads contribute some fraction to lithospheric load- 111 ing, T_e would be underestimated by considering 112 surface loading alone. Taking into account surface 113 and subsurface loading, Forsyth [1985] and a 114 number of subsequent studies recovered very large 115 T_e (>60 km) values for cratonic interiors. 116

[7] More recently, *McKenzie* [2003] suggested that 117 long-term erosion and sedimentation over cratons 118 might generate internal loads with no topographic 119 expression. These loads would constitute noise 120 within the flexural model framework, as they 121 would represent density anomalies with no appar- 122 ent flexural response. McKenzie [2003] argued that 123 this potential noise field would bias the spectral 124 methods results using the coherence function and 125 the Bouguer anomaly (Bouguer coherence) but not 126 those obtained from the admittance function and 127 the free-air gravity (free-air admittance). Using the 128 free-air admittance he obtained T_e estimates in 129 cratons that were <25 km. However, subsequently, 130 it has been shown that if the free-air admittance is 131 formulated consistently with the Bouguer coher- 132 ence, both methods yield similar results (within 133 uncertainties) when applied to synthetic as well as 134 real data (see Pérez-Gussinyé et al. [2004] and 135 Pérez-Gussinyé and Watts [2005] for a detailed 136 discussion). In Europe, for example, both techni- 137 ques yield low T_e in young Phanerozoic terranes 138 and large estimates in cratons ($T_e > 60$ km), where 139



the topography is subdued and long-term erosion and sedimentation have occurred. This suggests that subsurface loads without topographic expression may not occur [*Pérez-Gussinyé and Watts*, 2005]. Consequently, most spectral methods for T_e estimation use the Bouguer coherence function and a loading model that includes surface and subsur-

147 face loading, as is done here.

[8] That T_e may depend on the dimensions of the 148 data window stems from two modeling limitations. 149First, T_e is generally assumed to be constant within 150the data windows used, such that if the data 151encompasses regions with different rigidities, the 152estimated T_e is a weighted average. Secondly, when 153the data window is small relative to the flexural 154wavelength, the T_e estimate tends simultaneously 155to be biased toward lower values and have larger 156variance (see Swain and Kirby [2003], Audet and 157Mareschal [2004], Pérez-Gussinyé et al. [2004], 158and also section 4.1 for a detailed description). 159Additionally, techniques used for power spectral 160 estimation affect the wavelength content of the 161 topography and gravity data eventually yielding 162differences in the absolute values of T_e estimates 163(see Ojeda and Whitman [2002] and Audet and 164Mareschal [2004] for tests with differing spectral 165estimators and section 3.3 for a detailed description 166 of this effect). Despite this, relative spatial varia-167tions of T_e estimates tend to agree between differ-168ent studies such that the areas found to have 169highest or lowest T_e will persist for various 170methodologies. Yet, comparisons of T_e in different 171regions of a continent benefit from application of a 172single consistent estimation approach, as is carried 173out in this paper. 174

[9] In South America, the terrestrial Bouguer 175gravity anomaly data coverage is uneven and so 176most previous studies have been limited to regions 177where there is adequate data. Most have focused 178on the Andean domain [Tassara, 2005; Stewart 179and Watts, 1997; Watts et al., 1995]. An exception 180 are those of Mantovani et al. [2005a], who esti-181 mate T_e on the basis of an empirical correlation 182 between tidal forcing and elastic thickness. Another 183is a recent study by Tassara et al. [2007], who 184used wavelet Bouguer coherence to obtain spatial 185variations in T_e . The results of these two studies, 186 however, differ markedly: Mantovani et al. 187 [2005a] obtained large ($T_e > 70$ km) estimates 188 over the Andean domain comparable to those in 189cratonic South America. Tassara et al. [2007], on 190the other hand, found T_e in the Andes to be much 191

lower than over the cratonic interior. This latter 192 study is consistent with forward modeling results 193 of Bouguer anomalies in the Andean domain 194 [*Tassara*, 2005, *Stewart and Watts*, 1997; *Watts* 195 *et al.*, 1995]. 196

[10] In this paper, we present a new T_e map of 197 South America generated using the coherence 198 between continent-scale grids of Bouguer anomaly 199 and topography data together with a multitaper 200 power spectral estimator. We focus on the results 201 in the stable South American platform, first 202 describing the tectonic domains of cratonic South 203 America, the data sources and processing, and the 204 methodology employed to estimate T_e . We then 205 compare our results to those of Tassara et al. [2007] 206 and Mantovani et al. [2005a]. Subsequently, we 207 examine the relationship of T_e to terrane age and 208 other proxies for lithospheric structure such as 209 shear wave velocity and heat flow. Finally we dis- 210 cuss the spatial correlation of T_e with intraconti- 211 nental seismicity and its significance for neotectonic 212 deformation of the continental interior. 213

2. Tectonic Terranes of the South214American Platform215

[11] The South American stable platform com- 216 prises Archean and Proterozoic terranes amalgam- 217 ated during the Trans-Amazonian (Paleo-Proterozoic), 218 Late Meso-Proterozoic and the Brasiliano/Pan African 219 orogenies [*Almeida et al.*, 2000] (Figure 1). The super- 220 continents Atlantica, Rodinia and West Gondwana, 221 respectively, resulted from the culmination of these 222 three tectonic cycles [*Brito Neves et al.*, 1999]. 223

[12] Within the platform, the Amazonian, Sao 224 Francisco and Rio de la Plata cratons are the largest 225 cratonic blocks remnant after these cycles (Figure 1). 226 These blocks contain Archean nuclei as well as 227 fragments of Paleo-Proterozoic and Late Meso- 228 Proterozoic/Early Neo-Proterozoic fold belts [Brito 229 Neves et al., 1999]. The most extensive exposures 230 of Archean rocks are found within the Amazonian 231 and Sao Francisco cratons [Almeida et al., 2000]. 232 The Amazonian craton is divided into the Guyana 233 and Guaporé Shields, which consist of four differ- 234 ent geological terranes of Archean and Early/Middle 235 Proterozoic age. The Guyana shield is dissected 236 by a SW-NE trending graben, the Tacutu Meso- 237 zoic graben (Figure 1). The graben separates 238 blocks having different composition and age, so 239 probably it is a reactivated structure of Late/Middle- 240 Proterozoic age. 241





Figure 1. Topography of South America with main tectonic provinces within the stable platform. The Guaporé, Gua, and Guyana, Guy, shields have basement of the same nature and age ranging from Archean to Early/Middle Proterozoic. Together they form the Amazonian craton. Within the Guyana Shield, the Tacutu Mesozoic rift, TAC, is a failed rift branch associated with the opening of the central Atlantic in Lower Jurassic times. Late Triassic flood basalts belonging to the central Atlantic magmatic province (CAMP) are located at the northwestern end of the Tacutu graben (approximate location shown by dashed gray line [Phipps Morgan et al., 2004]), which separates blocks with different Archean and Proterozoic age suggesting an ancient suture that predates central Atlantic opening. TBL is the Transbrasiliano lineament, a continental scale suture that recorded the amalgamation of the Amazonian with the San Francisco, SF, and Rio de la Plata, RP, cratons during the Brasiliano/Pan African cycle. The location of the Rio de la Plata craton is poorly known, and we show it with a black dotted line. The Patagonian terrane, Pat, was amalgamated during the Phanerozoic. The red dot is the location of Asunción in Paraguay, and the gray lines represent dyke swarms that fed the Paraná flood basalts [Hawkesworth et al., 2000]. The extension of the flood basalts is shown with a black dashed line. Par is Paraná basin; Prb is Parnaiba basin. Both basins are thought to be underlain by cratonic lithosphere; see text. Am and Sol are Amazonian basin (east of 300°) and Solimoes basin (west of 300°), respectively. Within the Andean domain, iso-depth contours to the oceanic slab from Syracuse and Abers [2006] are shown with red dashed lines; numbers are depths in kilometers.



Figure 2. Figure 2a is the final Bouguer anomaly after merging the SAGAP data points shown in Figure 2b with the global model of 1° of resolution: EIGEN-CG30C, obtained from CHAMP and GRACE satellites and terrestrial data [*Foerste et al.*, 2005]. See text for a discussion on the procedure to obtain the Bouguer anomaly.

[13] The collision of the Amazonian craton to the 242north with the Sao Francisco and Rio de la Plata 243cratons to the south during the Pan-African/ 244Brasiliano orogeny is recorded by the Transbrasi-245liano megasuture (Figure 1) [Cordani and Sato, 2461999]. The suture is a continent-scale NE-SW ductile 247shear zone that extends into West Africa as the Hoggar 248suture [Trompette, 1994]. In South America the suture 249extends from northeast Brazil to the Bolivian border 250(Figure 1) [Trompette, 1994]. 251

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[14] The Rio de la Plata craton is thought to 252underlie Paleozoic and Mesozoic sedimentary 253successions beneath most of southeastern Brazil, 254Uruguay and northern Argentina [Trompette, 2551994; Basei et al., 2000]. The approximate loca-256tion of the craton is shown in Figure 1. The south-257ernmost part of South America is occupied by the 258Paleozoic Patagonia terrane. This terrane was 259accreted to the South American stable platform 260during the Hercynian orogeny, culminating in the 261amalgamation of Pangea [Ramos, 2000]. 262

[15] The South American stable platform includes 263a number of large Paleozoic basins: the Amazon, 264Solimoes, Parnaiba, and Paraná in Brazil, and the 265Chaco in Argentina. The Paraná basin in Brazil 266comprises Mesozoic basalts that were possibly 267derived from the Tristan da Cunha hot spot [Turner 268et al., 1994]. The dyke-swarms that fed the flood 269basalts have been mapped in Ponta Grossa in 270

Brazil and in a region eastward of Asunción in 271 Paraguay (Figure 1). Despite this magmatic epi- 272 sode, the basin is thought to be underlain by a 273 cratonic block on the basis of radiometric dating 274 [*Cordani et al.*, 1984], Bouguer anomaly studies 275 [*Mantovani et al.*, 2005b] and the high S-velocities 276 that have been imaged to 200 km depth [*Schimmel* 277 *et al.*, 2003; *Snoke and James*, 1997]. 278

3. Methodology 279

3.1. Generation of a Continent-Wide280Bouguer Anomaly Grid281

[16] To produce the continent-wide Bouguer anom- 282 aly shown in Figure 2, we have combined irregu- 283 larly distributed Bouguer gravity anomalies 284 compiled by GETECH as part of their South 285 America Gravity Project (SAGP) [Green and 286 Fairhead, 1991] with those derived from free-air 287 gravity anomaly data from the EIGEN-CG30C 288 model. To our knowledge, the compilation pre- 289 sented in this work constitutes the most accurate 290 gravity database used for T_e studies over the 291 whole South American continent. The EIGEN- 292 CG30C data set combines free-air gravity data 293 measured by the satellites CHAMP (860 days) and 294 GRACE (376 days) with various sources of ma- 295 rine and terrestrial free-air gravity including 296 EGM96 to yield a final grid of 1° lateral resolu- 297



Final Bouguer anomaly from GG30C and Getech data
 Getech data points



Figure 3. Tracks of the final Bouguer anomaly constructed using a spherical approximation to the Bouguer anomaly (red line) along three different latitudes: (a) 0° , (b) 20° S, and (c) 40° S. The flat slab approximation to the Bouguer anomaly (green line) and the Bouguer SAGAP data points (blue dots) are also shown. Note that in oceans the SAGAP data is free-air anomaly.

tion. The overall accuracy of the EIGEN-CG30C model at spatial scales of ~ 100 km is estimated to be 8 mgal [*Foerste et al.*, 2005].

[17] The coherence function used here to calculate 301 T_e , estimates the correlation of topography and 302 Bouguer gravity anomaly as a function of wave-303 length. Subsequently, the coherence is modeled to 304 determine the effective elastic thickness of the 305lithosphere. Therefore, to avoid introducing spuri-306 ous wavelengths in the Bouguer anomalies and 307 thereby noise in the coherence function, the topog-308 raphy used for the Bouguer correction must have 309 the same resolution as that of the free-air gravity 310 grid. Given that the EIGEN-CG30C free-air gravity 311 grid has a resolution of 1° and the length of a 312longitudinal degree varies by $\sim 60\%$ from northern 313 to southern South America, one cannot define a 314 wavelength in planar coordinates equivalent to 1° 315 of resolution over the entire continent. Therefore 316 we transformed a topography grid to spherical 317coordinates and calculated the Bouguer correction, 318

 $\Delta g(r, \theta, \phi)$ to 1° resolution following *Wieczorek* 319 and *Phillips* [1998] and *Lowry and Zhong* [2003]: 320

$$\Delta g(r, \theta, \phi) = \frac{(l+1)GM}{r} \sum_{ilm} \left(\frac{D}{r}\right)^{l+1} C_{ilm} Y_{ilm}(\theta, \phi)$$
$$C_{ilm} = \frac{4\pi\Delta\rho D^2}{M(2l+1)} \sum_{n=1}^{l+3} \left[\frac{{}^n h_{ilm}}{D^n n!} \frac{\prod_{j=1}^n (l+4-j)}{(l+3)}\right]$$

These equations yield the gravity anomaly, $\Delta g(r, \theta, 321 \phi)$, due to a topographic surface, $H(r, \theta, \phi)$, 323 referenced to a sphere of radius D, and are the 324 spherical equivalent to Parker's finite-amplitude 325 formulation of gravity due to topography on a 326 plane [*Parker*, 1972]. Here are the harmonic 327 coefficients corresponding to the spherical trans- 328 form of the nth power of the topography, M is mass 329 of the Earth and G is the universal gravitational 330 constant. 331

[18] Finally, the Bouguer anomaly derived from the 332 EIGEN-CG30C data was merged with the sparsely 333 distributed GETECH Bouguer anomaly to obtain 334 the final grid. The two data sets were combined 335 such that EIGEN-CG30C data were replaced by 336 GETECH data points where available. The final 337 grid spacing is 8 km, although the information 338 content generally corresponds to the 1° grid outside 339 of areas where dense GETECH data are available 340 (Figure 2). Before merging we tested for systematic 341 offsets between the two data sets. Figure 3 shows 342 the GETECH data (blue circles), the profiles of 343 the Bouguer anomaly constructed using the spher- 344 ical Bouguer correction (in red), and the profiles of 345 the Bouguer anomaly using a slab correction (in 346 green). The GETECH data points offshore are free- 347 air anomalies, so are systematically lower than our 348 final Bouguer anomaly offshore (Figure 3), but the 349 onshore anomalies are in close agreement. 350

[19] Our T_e analysis is implemented in Cartesian 351 coordinates, necessitating a projection of the spher- 352 ical data that minimizes distortion. We accom- 353 plished this by dividing South America into four 354 smaller grids north-to-south and projecting to pla- 355 nar coordinates within each area. We then back- 356 projected the Cartesian-coordinate T_e maps into 357 longitude and latitude for presentation. 358

3.2. Bouguer Coherence Using a Multitaper 360 Spectral Estimator 361

[20] The effective elastic thickness, T_e , of the 362 lithosphere is the thickness of an ideal elastic plate 363



that would bend by the same amount as the 364 lithosphere, under the same applied loads [e.g., 365 Watts, 2001]. Because layers composing the litho-366 sphere fail anelastically, the measured T_{e} is actually 367 an integral of the elastic bending stress, constrained 368 by the limits imposed by the brittle and ductile 369 rheologies of the lithosphere [Burov and Diament, 370 1995; Lowry and Smith, 1995]. 371

[21] To measure the elastic thickness, we use as 372 input data the topography (which sums surface 373 loads imposed on the lithosphere and the flexural 374 deflections that compensate both surface and sub-375surface loads), and the Bouguer anomaly (which 376 contains the mass signal from subsurface loads 377 plus the deflections caused by loading). The 378 coherence function between the topography and 379Bouguer anomaly, commonly known as Bouguer 380 coherence, gives information on the wavelength 381 band over which topography and Bouguer anom-382aly are correlated, and is given by 383

$$\gamma_{obs}^2 = \left\langle \frac{\left|S_{hb}(k)\right|^2}{S_{hh}(k)S_{bb}(k)} \right\rangle$$

where $S_{hb}(k)$, $S_{hh}(k)$, $S_{bb}(k)$ are the cross-power spectrum of the topography and Bouguer anomaly and the auto-power spectra of the topography and and the Bouguer anomaly, respectively. Angle brackets denote averaging over annular wave number bands of the wave number modulus $k = |k| = \sqrt{k_x^2 + k_y^2}$.

[22] The Bouguer coherence generally tends to 392zero at short wavelengths, where the topography 393 is not compensated, and it tends to one at long 394wavelengths where the response to loading is Airy-395like [Forsyth, 1985]. The wavelengths at which the 396 coherence increases from 0 to 1 depend on the 397 effective elastic thickness, T_e , of the lithosphere. 398 When the lithosphere is relatively weak and T_e is 399 small, local compensation for loading occurs at 400relatively shorter wavelengths. 401

[23] To estimate T_e , we compare the observed 402coherence with the coherence curves predicted 403 for a particular set of T_e values. The T_e that 404minimizes the difference between the predicted 405and observed coherence is the assigned T_e for an 406 analyzed area. To calculate the predicted coher-407ence, assumptions about the loading processes in 408the lithosphere need to be made. We follow 409Forsyth [1985] and assume that surface loads 410(atop the lithosphere) and subsurface loads (within 411

the lithosphere) are statistically uncorrelated. Sur- 412 face loads include the thrust sheets that comprise 413 topography in orogenic belts while subsurface 414 loads include intracrustal thrusts and magmatic 415 underplating. For any given T_e , we calculate a 416 set of surface and sub-surface loads and compen- 417 sating deflections that reproduce exactly the ob- 418 served topography and gravity anomaly, an 419 approach commonly known as load deconvolution 420 [*Forsyth*, 1985]. Using this approach, the ratio of 421 surface to subsurface loads, or loading ratio, varies 422 with two-dimensional wave number and is not 423 imposed as an independent parameter as when 424 analytical solutions are calculated. 425

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[24] Forsyth's [1985] original formulation of the 426 predicted coherence assumes that all internal den- 427 sity variation and loading occurs at the Moho. We 428 used CRUST2.0 [Bassin et al., 2000] to define the 429 internal density profile and assumed that internal 430 loading occurs at the interface between upper and 431 mid-crust. The lateral variation in depth of this 432 interface was obtained from CRUST2.0. Since the 433 observed coherence can be reproduced equally well 434 by either low T_e and shallow loading or a larger T_e 435 and deeper loading, there is a trade-off between T_e 436 and assumed depth of loading. However, we tested 437 the sensitivity of Te to loading depth in Europe and 438 found that changing the loading depth from the 439 mid-crust to Moho changed T_e by ~5 km, but the 440 general patterns of variations remained the same 441 [Pérez-Gussinvé and Watts, 2005]. 442

[25] Although we do not explicitly include the 443 subducting Pacific slab in our loading model, 444 modeling of the slab Bouguer anomaly signal 445 expected for the ≥ 100 km depth contours shown 446 in Figure 1 indicates that signal is dominated by 447 wavelengths that are much longer than the wave- 448 lengths of flexural transition or even the window 449 dimensions used in this analysis. The subducting 450 slab correlates with the ~ 2000 km long-wave- 451 length anomalies of the Bouguer anomaly field 452 (but has opposite sign, such that main effect of 453 the slab is to offset the gravity signal associated 454 with thickened crust in the Andean plateau by 455 5-10%). The longest wavelength used for anal- 456 vsis here is <800 km, corresponding to the 457 largest window dimension. Hence, at wave- 458 lengths of the flexural transition, slab dynamical 459 effects in gravity and topography change the 460 coherence of the two negligibly. Re-estimation 461 of T_e after subtracting the estimated slab signal 462 from the Bouguer data yields a negligible change 463 in the estimates [Pérez-Gussinyé et al., 2006]. 464





466 3.3. Resolution Tests With Synthetic Data

[26] Calculation of the observed and predicted 467coherence involves transformation into the Fourier 468 domain of the topography and Bouguer gravity 469anomaly to estimate their auto- and cross-power 470471spectra. Because both data sets are non-periodic and finite, the Fourier transformation presents 472problems of leakage, or transference of power 473between neighboring frequencies, resulting in esti-474mated spectra that differ from the true spectra. To 475reduce leakage, the data are tapered prior to Fourier 476transformation. However, ultimately, the type of 477 taper used influences slightly the resulting power 478spectra and hence the coherence function. Hence 479the ability to recover T_e differs depending on the 480 tapering technique used, making it important to 481understand its limitations. 482

[27] In this paper, we use Thomson's multitaper 483method [Thompson, 1982] with Slepian windows 484 [Slepian, 1978]. The spectral estimator obtained 485 with the multitaper is a weighted average of the 486spectra generated with a set of individual, orthog-487onal tapers. The multitaper estimator reduces the 488variance of the spectral estimate and also defines 489spectral resolution [Percival and Walden, 1993]. 490 The set of orthogonal tapers are defined by setting 491 the bandwidth of the central lobe of the power 492spectral density of the first-order taper, W. For a 493given W, there are a maximum of K = 2NW - 1494number of tapers, with good leakage properties that 495can be used for the estimation of the spectra, where 496 N is the number of samples within the data window 497[e.g., Percival and Walden, 1993; Simons et al., 498 2000]. Variance of the spectral estimates decreases 499with the number of tapers used as 1/K, so that the 500bandwidth and number of tapers are chosen 501depending on the function under analysis [Percival 502and Walden, 1993]. We use here a multitaper 503scheme corresponding to NW = 3, which is also 504used in many other studies for T_e estimation [e.g., 505Audet and Mareschal, 2004; Pérez-Gussinyé et al., 5062004; Pérez-Gussinvé and Watts, 2005]. Addition-507ally, we deconvolve the loads within the same data 508window as was used to derive the observed coher-509ence. Pérez-Gussinvé et al. [2004] deconvolved 510loads in a window larger than that used to calculate 511the observed coherence/admittance. Subsequently, 512the loads resulting from the deconvolution were 513transformed back into the spatial domain and 514windowed within the same region and multitaper 515parameters as the observed coherence/admittance. 516Here we have deconvolved the loads and calculated 517518the power spectra within the same window to speed up the calculation of T_e . We show here tests with 519 synthetic data of the method's ability to recover T_e 520 using this slight variation on the method. These 521 tests indicate that estimation bias and variance 522 similar to those given by *Pérez-Gussinyé et al.* 523 [2004] can be achieved using smaller T_e estimation 524 windows. 525

[28] The generation of synthetic data has been 526 explained in detail by Pérez-Gussinvé et al. 527 [2004] and will only be briefly described here. 528 The synthetic topography and gravity data was 529 generated by placing uncorrelated surface and 530 subsurface mass loads on an elastic plate using 531 an algorithm similar to that of Macario et al. [1995]. 532 First, the Fourier amplitudes of uncorrelated sur- 533 face, Hi, (k), and subsurface, Wi, (k), mass loads 534 were calculated imposing that their spectra follows 535 a power law distribution with respect to wave num- 536 ber, with a fractal dimension of 2.5, as it is observed 537 for the amplitude spectra of the real topography 538 [Mandelbrot, 1983; Turcotte, 1997]. Surface and 539 subsurface loads were then standardized (unit vari- 540 ance) and their amplitudes were scaled so that their 541 loading ratio has an expected value of 1, although 542 the loading ratio varies with wave number. Vertical 543 stresses $ho_{\rm c}$ gHi and $(
ho_{\rm m}ho_{\rm c})$ gWi, where g is 544 gravitational acceleration and $\rho_{\rm c}$ and $\rho_{\rm m}$ are the 545 densities of the crust and mantle respectively, were 546 applied as loads at the surface and Moho, respec- 547 tively, of a thin elastic plate with a specified elastic 548 thickness T_e . In the case where T_e is spatially 549 constant, the amplitudes of the final topography, 550 H, and Moho deflection, W, were calculated from 551 the load response relations given in Appendix A of 552 Pérez-Gussinyé et al. [2004]. The Bouguer anomaly 553 was calculated from the Moho deflection using the 554 Parker [1972] finite amplitude formulation up to a 555 fourth-order approximation. We computed 100 such 556 sets of topography and Bouguer anomaly by chang- 557 ing the random generator seed. 558

[29] In Figure 4 we show tests of synthetic topog- 559 raphy and Bouguer anomaly generated with a 560 spatially constant *Te* and multitaper parameters of 561 NW = 3 and K = 5 (as done by *Pérez-Gussinyé et* 562 *al.* [2004]). The grid interval is 8 km. The tests 563 show that when the flexural wavelength, $\lambda = \pi (4D/564 \Delta \rho g)^{1/4} \sim 29 T_e^{3/4}$ [*Swain and Kirby*, 2003], is 565 large relative to the window size, some of the 566 resulting *Te* values are underestimated and that 567 additionally, the number of spuriously high *Te* 568 estimates, or outliers, increases (as previously 569 observed by *Swain and Kirby* [2003], *Audet and* 570 *Mareschal* [2004], and *Pérez-Gussinyé et al.* 571





Figure 4. Tests with synthetic data and constant T_e for (top) 800 × 800 km windows and (bottom) 400 × 400 km windows. The synthetic data used are the same as those used by *Pérez-Gussinyé et al.* [2004], but the tests shown here used smaller windows and a different load deconvolution routine (see section 3.3). The legend details the input T_e value, the mean output T_e , standard deviation, and number of outliers for 100 tests. Note that for a window of a given size, the number of outliers (defined as $T_e > 120$ km) increases as the true T_e increases.

572 [2004]). Therefore, when T_e is constant, larger 573 windows produce more accurate estimates.

[30] However, T_e is likely to vary in different 574575geological terranes. To generate synthetic data with spatially varying T_e , we transformed the initial 576surface and subsurface loads to the spatial domain 577 and solved the fourth-order flexural governing 578equation using a finite difference solution [Wyer, 5792003; Stewart, 1998]. In order to retrieve a spa-580tially varying T_e structure, the T_e analysis was 581carried out using constant sized, overlapping win-582dows with centers spaced 56 km apart. Within each 583window T_e was assumed to be constant, and the T_e 584estimate was assigned to its center. 585

⁵⁸⁶ [31] Figure 5 shows estimates from synthetic data generated with the spatially varying structure shown in Figure 5d, using multitaper parameters NW = 3 and K = 5 and three different windows of 400×400 km, 600×600 km and 800×800 km (see *Pérez-Gussinyé et al.* [2004] for results with larger windows). For the smallest window, the values toward the centre are overestimated and 593 those around the high T_e nuclei are underestimated 594 (Figure 5c). The 600×600 km window recovers 595 the spatial variations better although it still over- 596 estimates the highest T_e values (Figure 5b). Finally, 597 the larger window size recovers the highest 598 T_e values better but overestimates the lower T_e 599 values due to spatial averaging (Figure 5a). Hence 600 there is a trade-off between spatial resolution, 601 which ideally would be better with smaller win- 602 dows, and the ability to recover large T_e values, 603 which should improve with larger windows [see 604 also Pérez-Gussinyé et al., 2004]. Here we choose 605 to use a Fourier windowing technique based on the 606 multitaper and analyze the T_e in South America 607 using 3 different window sizes. 608

[32] To estimate T_e in South America we did 609 several tests with different NW values and number 610 of tapers. The pattern of T_e variation is similar with 611 the different multitaper parameters used, but the 612 mean values change, as they do when the window 613 size changes (see Figure 5). Here, we present the 614 results for NW = 3 and 5 tapers for all of South 615



Figure 5. Recovery of the variable T_e structure shown in Figure 5d, using windows of different sizes. The synthetic data are generated as those used by *Pérez-Gussinyé et al.* [2004]. (a) 800 × 800 km, (b) 600 × 600 km, and (c) 400 × 400 km. Multitaper parameters are NW = 3 and K = 5 (see section 3.3 for the meaning of these parameters). Gray circles in Figures 5a, 5b, and 5c represent the contoured T_e values shown in Figure 5d. The smallest windows, Figure 5c, tend to overestimate the largest T_e values (50 km) and underestimate intermediate T_e values ranging from 20 to 30 km. The largest windows, Figure 5a, approximate the largest T_e better but overestimate $T_e < 20$ km due to spatial smearing and averaging (see text for a detailed description and also *Pérez-Gussinyé et al.* [2004] for similar tests with larger windows).

America (Figure 6). We have chosen these param-616 eters because the larger number of tapers gives 617 a relatively smooth solution over stable South 618 America. In this area the topography has low relief, 619 so that errors and dynamical noise fields in the 620 topography and gravity data can have relatively 621 greater effect on the coherence of the two data sets. 622Consequently, it is important to have stable (small 623 variance) power spectral estimates, which implies 624 625 the use of 5 tapers, the maximum allowed for NW = 3 (note that the variance in the spectra 626 decreases with 1 over the number of tapers). We 627 have performed tests with 3 tapers and small 628 windows (400 \times 400 km) over the stable platform, 629 and found that this combination of taper parameters 630 and window size is unstable, as it produces large 631 variance in the coherence estimates, making it 632 difficult to differentiate between real geological 633structure and noise in the T_e estimates. Therefore 634 we use a larger number of tapers, to obtain a smooth 635 T_e structure even with small windows, which allows 636 having a detailed lateral resolution of the structures 637 within the stable platform (Figure 6a). 638

640 **4. Results**

641 [33] We have estimated T_e using windows of 400 × 642 400 km, 600 × 600 km and 800 × 800 km. Since 643 large T_e values cannot be recovered with confidence using such small windows, we plot $T_e > 644$ 70 km in black in Figure 6. Our results show a 645 first-order T_e variation in South America in which 646 the Andes have relatively low T_e and the stable 647 platform has relatively high values (Figure 6). 648 Although the first-order pattern of spatial variation 649 in T_e remains similar for the three estimation 650 windows, there are differences in the lateral extent 651 of the imaged structures as well as the mean T_e 652 recovered. These differences are analogous to those 653 observed when synthetic data are used (Figures 4 654 and 5). For example, within the predominantly 655 high T_e stable platform, areas of low T_e are more 656 prominent with the smallest windows, as these 657 image lateral boundaries in T_e better. This is 658 because larger windows spatially average small 659 areas with low T_e embedded within areas with 660 large T_e such that they may be damped or disappear 661 entirely. For example, compare the relatively low 662 T_e area along the Transbrasiliano lineament in 663 Figures 6a and 6c. 664

[34] Along the Andean orogenic belt and flanking 665 foreland, T_e also changes considerably with win- 666 dow size. For example, along the Andean foreland 667 there is an area of high T_e , centered at ~65°W and 668 18°S, which has an arcuate shape and is clearly 669 distinguished in the 800 × 800 km and 600 × 670 600 km window results (Figures 6b and 6c) but is 671 Geochemistry Geophysics Geosystems



Figure 6





nearly invisible in the 400 \times 400 km estimates 672 (Figure 6a). We interpret that the high T_e observed 673 in the 800 \times 800 km and 600 \times 600 km windows 674 reflects a relatively strong part of the Brazilian 675 Shield that underthrusts the sub-Andean fold and 676 thrust belt. Our interpretation is based, in part, on 677 previous 1-D analysis along the Andean foreland in 678 the bend region which observed a similar structure 679 [Stewart and Watts, 1997; Watts et al., 1995]. Local 680 tomography in the vicinity of 16°S also exhibits a 681 strong lateral change in P wave velocity beneath 682 the Eastern Cordillera ($\sim 65.5^{\circ}$ W), with high ve-683 locities to the east interpreted as underthrusting 684craton [Dorbath et al., 1993]. Shear wave anisot-685 ropy studies also show a change of the fast prop-686 agation direction from north-south to east-west, 687 688 consistent flow patterns expected if the edge of the cratonic lithosphere is at $\sim 65.5^{\circ}$ W [Polet et 689 al., 2000]. Thus we believe that the 400 \times 400 km 690 window results miss part of the actual T_e structure 691 in this area. Two localized patches of very high 692 693 T_e (>70 km) are observed in the small window estimates for this area. We interpret that these high 694 rigidities occur because the flexural wavelength is 695large in relation to the window, and the rest of the 696 T_e values are underestimated here similar to what 697 we observe in Figures 4 and 5. 698

[35] Despite the differences in estimates of T_e along 699 the Andean chain, some relative T_e variations are 700 common to all three estimation windows. These 701 include: increasing rigidity from 40° S to $\sim 20^{\circ}$ S in 702 the region overlying the 50 to ~ 100 km depth 703 contours to the subducting slab, and comparatively 704 lower rigidity in continental lithosphere overlying 705 the subducted Nazca and Carnegie ridges (Figure 6). 706 In subsequent work we detail window and multi-707 taper estimation parameters, which best describe 708 the lateral variation of T_e in the Andean domain. 709

There, the significance of these variations for the 710 thermal and mechanical structure and the geody-711 namics of the Andean domain will be discussed 712 (M. Pérez-Gussinyé et al., Spatial variations in the 713 effective elastic thickness, T_e , along the Andes: 714 Implications for subduction geometry, manuscript 715 in preparation, 2007; hereinafter referred to as 716 Pérez-Gussinyé et al., manuscript in preparation, 717 2007). In the next section we compare our results 718 with previous studies. Subsequently, we discuss the 719 relationship of T_e to the tectonic, seismic and 720 thermal structure of the stable platform as well as 721 its significance for neotectonic deformation of the 722 continental interior.

4.1. Comparison With Previous Results 724

[36] Previous estimates of T_e for the entire conti- 725 nent of South America include those of Mantovani 726 et al. [2001, 2005a] and Tassara et al. [2007]. Our 727 results are in agreement with the findings of 728 Tassara et al. [2007] and with forward modeling 729 of Bouguer anomalies along profiles of the Andean 730 domain [Tassara, 2005; Stewart and Watts, 1997; 731 *Watts et al.*, 1995], all of which indicate that T_e in 732 the Andes is much lower than over the cratonic 733 interior. However, they differ markedly from those 734 of Mantovani et al. [2001, 2005a], which indicate 735 large T_e (~70-80 km) over the Andean domain, 736 comparable to estimates in cratonic South America. 737 Mantovani et al. [2001, 2005a] estimated T_e using 738 an empirical correlation between tidal gravity 739 anomalies and elastic thickness, and used this 740 correlation to generate elastic thickness values 741 where terrestrial gravity measurements where 742 sparse. Given that tidal loading is of short duration 743 (about one day), and the lithospheric deformation 744 in response to such short-term cyclical stress is 745 quite different than to longer-term geological pro- 746

Figure 6. (top) T_e estimates for South America for three different window sizes, (a) 400 × 400 km, (b) 600 × 600 km, and (c) 800 × 800 km, and multitaper parameters of NW = 3 and 5 tapers. (Note that black colors indicate indeterminately large T_e). (bottom) The same as in top, but T_e is superimposed by a normalized catalogue of earthquakes within Brazil, Paraguay, and Uruguay (see *Assumpçao et al.* [2004] for a description of the normalization), by the depths to subducted slab (50 to 250 km from *Syracuse and Abers* [2006]) and by the main tectonic provinces. (d) Bathymetry of South America offshore, and the age of igneous and metamorphic rocks believed to indicate the age of crustal formation [*Schobbenhaus and Bellizia*, 2001]. These are overlain by the main tectonic provinces and the depths to the slab [*Syracuse and Abers*, 2006]. Abbreviations are as in Figure 1, except for Chc, which is Chaco basin. (e) Heat flow anomaly which results from subtracting a regional heat flow field from the observed heat flow values [*Hamza et al.*, 2005]. Triangles are heat flow measurements. The regional heat flow field is a polynomial representation of the South American heat flow and is meant to represent the first-order increase of 60 mW/m² in the Stable Platform to 70 mW/m² in the Andes [*Hamza et al.*, 2005]. The heat flow anomaly is superimposed by the depths to the slab [*Syracuse and Abers*, 2006]. (f) Shear wave velocity at 100 km depth (from Feng et al., submitted manuscript, 2007) superimposed by the tectonic provinces, the seismicity from the normalized catalogue, and the depths to the slab [*Syracuse and Abers*, 2006].



cesses modeled here [Willett et al., 1984, 1985], the 747 T_e estimates of Mantovani et al. [2001, 2005a] are 748 unlikely to be comparable to those estimated here. 749Moreover, the correlation analysis of tidal gravity 750anomalies and T_e used by Mantovani et al. [2001, 751 2005a] utilized the very few South American T_e 752measurements that were available at that time, and 753 this may have biased their results. 754

[37] The study of Tassara et al. [2007] used 755Bouguer coherence to estimate T_e , and thus their 756results are directly comparable to ours. One differ-757 ence between their analysis and ours is that their 758Bouguer gravity anomaly was derived from the 759EIGEN-CG30C data set, which was combined in 760this work with the terrestrial GETECH data to 761 obtain a more detailed data set. An additional 762 important difference is that they used wavelets to 763 find the spatial variations in T_e , while we used a 764 Fourier windowing scheme with the multitaper 765method. When estimating spatially varying T_e 766 using a data-windowing approach, we are faced 767 with the dilemma that the small windows that 768would best resolve spatial variability in T_e cannot 769 recover large elastic thicknesses. However, larger 770 windows that adequately recover high T_e values 771 tend to smooth out lateral contrasts in T_e , yielding a 772 spatially smoothed version of the underlying T_{e} 773 distribution. The window size restricts the largest 774 T_e that can be recovered with confidence, in our 775case ~ 70 km. 776

[38] In this study, we circumnavigate the problem 777 of choosing an optimal window by presenting 778 results for different window sizes and interpreting 779 them according to the results obtained with syn-780 thetic data. Tassara et al. [2007] chose the alter-781 native approach of using a wavelet transform to 782 estimate a local coherence function at each point 783 of the data grid (an approach originally developed 784by Kirby and Swain [2004], Kirby [2005], and 785 Swain and Kirby [2006]). To perform the wavelet 786 transform, the signal (e.g., topography or gravity 787 anomaly) is convolved with a family of wavelets, 788 which have a range of dimensions, or scales (for a 789 more detailed description, see Kirby and Swain 790[2004]). Small-scale wavelets reveal the short-791 wavelength information content of the data, while 792large-scale wavelets reflect long-wavelength infor-793 mation [Swain and Kirby, 2006]. Convolution 794 effectively yields a wavelet transform estimate at 795every node of the data grid, enabling construction 796 of wavelet cross- and auto-power scalograms (the 797 wavelet equivalent to the Fourier domain cross-798and auto-power spectra), and hence the wavelet 799

coherence, at each data node and for each wavelet 800 scale [*Kirby and Swain*, 2004]. Because scale can 801 be mapped to wave number, the wavelet coherence 802 can be represented also as a function of the latter 803 [*Kirby and Swain*, 2004]. 804

[39] Thus, ideally, the wavelet transform method 805 would solve the dilemma of the choice of window 806 size by essentially using a different window for 807 each wave number and location [Kirby and Swain, 808 2004]. However, the T_e estimates of Tassara et al. 809 [2007] have a much smoother appearance than 810 those presented in this paper. For example, the 811 results presented here and those of Tassara et al. 812 [2007] both show the southwestern part of the 813 Transbrasiliano lineament to have low T_e (compare 814 Figures 6a and 6b with Figure 3 of Tassara et al. 815 [2007]). In our results, the low T_e area continues 816 along trend to the north, while the wavelet method 817 does not resolve this northward continuation, sim- 818 ilar to what occurs when we use large windows 819 (Figure 6c). Likewise, the area west of the Paraná 820 flood basalts, the Amazon basin and the Tacutu 821 graben appear as clear lineaments of low T_e in our 822 results (Figure 6a and partly Figure 6b) but are not 823 distinguished in the work of Tassara et al. [2007]. 824

[40] Hence, despite similarities in the loading mod- 825 els used, our results and those of Tassara et al. 826 [2007] differ markedly in the resolution of the low 827 T_e areas within the stable continent, raising the 828 obvious question of why. Differences between our 829 results and those of Tassara et al. [2007] could be 830 attributed to some combination of (1) the greater 831 information content of the gravity anomaly data 832 used in our analysis, (2) greater variance of multi- 833 taper windowed T_e estimates when the window is 834 not much larger than the wavelength of flexural 835 transition, and (3) reduction of information content 836 in the wavelet analysis of Tassara et al. [2007] by 837 inverse-wave number weighting of the norm of 838 predicted minus observed coherence. 839

[41] In order to test whether the greater accuracy of 840 the Bouguer anomaly data used here relative to that 841 used by *Tassara et al.* [2007] could explain the 842 differences in the results, we have re-computed T_e 843 using only the EIGEN-CG30C. Figure 7 shows 844 that the low T_e area in the Amazonian basin is less 845 pronounced using only EIGEN-CG30C data. How- 846 ever, other low T_e areas within the continental 847 interior are very similar in extent using both data 848 sets. This suggests that the greater information 849 content of the Bouguer data used here cannot by 850 itself account for the difference in the results. 851





Figure 7. T_e of South America estimated using 400 × 400 km windows and (a) the Bouguer gravity anomaly derived only from the EIGEN-CG30C data. (b) The same as in Figure 7 but using the data generated for this study (see section 3.1). These data result from combining the EIGEN-CG30C-derived Bouguer anomaly and the more detailed SAGAP Bouguer anomaly.

[42] Alternatively, one might interpret that the 852 windowed multitaper spectral approach has greater 853 variance and that many of the low T_e patterns in 854 Figure 6 are spurious. We deem this unlikely, 855 however, given the correlation of low T_e with 856 intracontinental seismicity and regions of greater 857 geological strain (see section 5). Similar correla-858 tions have been observed elsewhere [e.g., Lowry 859 and Smith, 1995] and are consistent with model 860 predictions of deformation focusing in regions of 861 thin lithosphere and/or low-viscosity upper mantle 862 [e.g., Latychev et al., 2005]. This coupled with the 863 fact that many of these features are persistent, 864 albeit attenuated, in estimates using all three win-865 dow sizes suggests that they represent real varia-866 tions in lithospheric flexural rigidity. Indeed, the 867 Tassara et al. [2007] estimates are most similar to 868 those we attain when using the largest window size 869 (Figure 6c), which yields a spatially smoothed 870 version of the weak T_e lineaments observed within 871 the cratonic interior (Figure 6a). 872

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[43] The wavelet method used by *Swain and Kirby* 873 [2006] and Tassara et al. [2007] weighted the 874 misfit between predicted and observed coherence 875 functions by the inverse wave number in order to 876 downweight the spurious high coherence at large 877 wave numbers that can result from an unrepresen-878 tative Bouguer reduction density or from certain 879 algorithms for generating gravity anomalies. On 880 the other hand, the wavelet approach includes a 881

range of wavelet scales that significantly exceeds 882 the wavelength of flexural transition. Including 883 large-scale wavelets in the coherence analysis is 884 somewhat analogous to using large windows in a 885 windowed spectral approach, and while the prac-886 tice of upweighting the longest wavelengths does 887 reduce the incidence of spurious low T_e estimates, 888 the greater emphasis placed on information 889 contained in the longest wavelengths of the data 890 also tends to smooth fine-scale structure (C. Swain, 891 personal communication, 2007). 892

[44] Despite their differences, multitaper and wave- 893 let estimates can be complementary if their respec- 894 tive limitations are considered when interpreting 895 the results. While we believe that we attain higher 896 spatial resolution of the low T_e features within the 897 rigid cratonic interior, the *Tassara et al.* [2007] 898 estimates better resolve T_e variations when $T_e > 899$ 70 km, allowing them to assess which regions have 900 the greatest rigidity within the continent. In future, 901 we plan a more rigorous comparison of the meth- 902 ods using synthetic data. 903

5. Discussion

5.1. Te and Tectonic Provinces

[45] Within the stable platform, T_e values are 907 generally high (>70 km). The figures show that 908 high T_e prevails over most of the Guaporé, Guyana 909

905

906



and Sao Francisco cratons (Figure 6), which largely 910 consist of Archean to Paleoproterozoic basement 911 [Trompette, 1994]. High T_e is also found in the 912Parnaiba and Paraná basins. Radiometric ages of 913 basement samples and the geometry of the sur-914 rounding fold belts indicate these basins to be cored 915 by very old (> 2 Ga) basement [Cordani et al., 916 1984; Brito Neves et al., 1984]. Relatively large T_e 917 has also been estimated from Bouguer coherence by 918 Vidotti [1997] and Vidotti et al. [1998] in the Paraná 919and Parnaiba basins using regional gravity anomaly 920 921 grids. Large T_e is also found in the approximate area of the Rio de la Plata craton (Figure 6). Most of the 922 Rio de la Plata craton is overlain by younger 923 sedimentary sequences of the southern Paraná and 924southwestern Chaco basins, where its basement is of 925926 unknown age, but the few exposures that are found are Archean in age [Trompette, 1994]. 927

[46] In the southern part of the stable continent, 928 south of the Patagonian suture, T_e is generally low 929 and the basement is Paleozoic in age (Figure 6). 930 Thus, in South America, the rigidity of terranes 931 which are mainly Archean to Paleo/Middle-932 Proterozoic appears, in general, to be much larger 933 than that of Phanerozoic terranes, which is consis-934tent with observations made on the European 935 lithosphere [Pérez-Gussinyé and Watts, 2005]. 936 Conductive cooling alone cannot explain the dif-937 ference in Te between these terranes, as the Paleo-938 zoic Patagonian terrane has had more than sufficient 939 time to conductively cool and stabilize. We suggest 940 that this contrast in T_e between old (>1.5 Ga) and 941 younger terranes probably reflects the larger thick-942 ness, depletion and, importantly, smaller water con-943 tent of old (> \sim 1.5 Ga) lithosphere, as has been 944 suggested for Europe [Pérez-Gussinyé and Watts, 9452005] and the western United States [Lowry and 946 Smith, 1994]. 947

[47] Within the Precambrian basement, there are 948four areas of relatively low T_e that coincide with 949sutures, rifts and sites of hot spot magmatism. 950 Starting from the south, we encounter low T_e in 951the Paraná basin at $\sim 27^{\circ}$ S, located west of the 952 region covered by flood basalts, where NW-SE 953 oriented dyke swarms that fed the \sim 130 Myr Paraná 954flood basalts are exposed (Figure 6) [Hawkesworth 955et al., 2000]. The large area of low T_e reaffirms the 956 large scale of the feature evident in aeromagnetic 957 surveys [Milner et al., 1995]. Further west, this low 958 T_e appears to propagate into the Chaco basin for the 959 smallest window size used here, which also has the 960 highest lateral resolution (Figure 6a). 961

[48] To the north, low T_e is observed along the 962 SW-NE oriented Transbrasiliano suture from $\sim 24^{\circ}$ 963 to $\sim 14^{\circ}$ S. This suture resulted from the collision of 964 the Amazonian with the Sao Francisco and Rio de la 965 Plata cratons during the Pan African orogeny 966 [*Cordani and Sato*, 1999]. The southwestern part 967 of this suture accommodated several igneous prov- 968 inces consisting of alkaline intrusions at $\sim 85-$ 969 60 Ma. These have been related to the Trindade 970 hot spot, currently located offshore Brazil in the 971 Trindade islands [*Gibson et al.*, 1995, 1997]. 972

[49] Along the eastern side of the Amazon basin, 973 we estimate a low T_e region which, for the smallest 974 window size used here, appears to continue west- 975 ward into the Solimoes basin (Figure 6a). This low 976 T_e may reflect Phanerozoic rifting that separated 977 the Brazilian craton into the Guyana and Guaporé 978 Shields. 979

[50] Finally, within the Guyana shield there is a 980 linear SW-NE trending zone of low T_e along the 981 Mesozoic Tacutu graben (Figure 6a). Coherence 982 analysis centered in this area similarly finds low T_e 983 [Ojeda and Whitman, 2002]. This graben likely 984 formed along an older existing suture because it 985 separates geological units of different Archean 986 and Early/Middle Proterozoic age and nature 987 [Trompette, 1994]. Moreover, this old suture may 988 have been reactivated during the Late Triassic, 989 providing the conduit through which flood basalts 990 of the Central Atlantic Magmatic Province were 991 emplaced at its northwestern end in Venezuela 992 (Figure 6). Later tectonism along this suture 993 occurred in Early Jurassic, when it functioned as 994 a rift branch of the early stages of opening of the 995 central Atlantic [Sears et al., 2005]. In summary, 996 within cratonic South America, low T_e often 997 coincides with locations where tectonism has 998 occurred repeatedly through geologic time. 999

5.2. Te and S-Wave Velocity

1001

[51] Shear wave velocities are sensitive to rock 1002 temperature and composition, which also influence 1003 rheology and hence the integrated strength of the 1004 lithosphere. Consequently we expect a positive 1005 correlation between S velocity and T_e . Comparison 1006 of T_e and S velocity is not straightforward as, 1007 analogously to T_e estimates, shear wave velocity 1008 models depend on the methodology employed for 1009 their estimation and the data used for analysis. 1010 Moreover, the relative sensitivities of T_e and 1011 S velocity to temperature, mineral composition, 1012 fluids and partial melt are different. Nevertheless, 1013





1014 the latest models of S velocity for South America 1015 [*Heintz et al.*, 2005; *Feng et al.*, 2004; M. Feng et 1016 al., Upper mantle structure of South America from 1017 joint inversion of waveforms and fundamental-1018 mode group velocities of Rayleigh waves, submitted 1019 to *Journal of Geophysical Research*, 2007 (herein-1020 after referred to as Feng et al., submitted manuscript, 1021 2007)] compare favorably to our results. For visual 1022 comparison we show in Figure 6 the S-velocity 1023 model of Feng et al. (submitted manuscript, 2007) 1024 at 100 km depth. Their shear wave velocities are 1025 obtained from simultaneously inverting regional S 1026 and Rayleigh waveforms and fundamental-mode 1027 Rayleigh-wave group velocities (Feng et al., 1028 submitted manuscript, 2007).

1029 [52] The resulting shear wave velocity and T_e maps 1030 correlate, in general, remarkably well. Common 1031 features found in the various tomographic models 1032 include high S velocity under the Guaporé, Guyana 1033 and Sao Francisco cratons and the Paraná and 1034 Parnaiba basins. All of these areas also exhibit 1035 high T_e (Figure 6). Note that T_e in excess of 70 km 1036 cannot be estimated with the window sizes used in 1037 our analysis, so we do not resolve differing T_e 1038 within regions of cratonic lithosphere. Hence, 1039 while S tomography indicates differing thicknesses 1040 for the cratonic Guaporé and Guyana shields, we 1041 depict these cratons as having apparently uniform 1042 high T_e (>70 km).

1043 [53] On the other hand, low S velocity has been 1044 inferred under the southwestern end of the Trans-1045 brasiliano lineament (TBL) [*Feng et al.*, 2004; 1046 *Heintz et al.*, 2005; Feng et al., submitted manu-1047 script, 2007] consistent with low P velocities in the 1048 SW and central segments of the TBL derived in 1049 local tomography with greater lateral resolution 1050 [*Schimmel et al.*, 2003; *Assumpçao et al.*, 2004]. 1051 The area of low P velocity along the TBL from the 1052 latter studies corresponds remarkably well with the 1053 low T_e estimated with the smallest window size 1054 (Figure 6a).

1055 [54] Along the Amazonian-Solimoes basin, low 1056 S-velocities have been imaged by *Heintz et al.* 1057 [2005] but were not observed across the entire 1058 basin by *Feng et al.* [2004] and Feng et al. 1059 (submitted manuscript, 2007). Within the Amazon 1060 basin, Feng et al. (submitted manuscript, 2007) 1061 obtained lower S-wave velocities than in the sur-1062 rounding shields (Figure 6f), which coincide with 1063 low T_e areas (Figures 6a and 6b), further supporting 1064 the view that the Amazon rifting system affected at 1065 least some parts of the lithosphere. [55] Along the Andean plateau there is good cor- 1066 relation between slow seismic velocity and low T_e . 1067 The relationship is particularly notable where aseis- 1068 mic ridges subduct (Figures 6b and 6f). In areas 1069 where flat subduction occurs, i.e., southern Chile, 1070 northern Peru and perhaps northern Colombia, 1071 S velocity (Figure 6f) and T_e are, in general, 1072 relatively larger. For example, relatively greater 1073 T_{e} is clearly observed in northern Colombia with 1074 all three estimation windows (Figures 6a, 6b, 1075 and 6c). Greater T_e is also observed in northern 1076 Peru but is most clear with 600×600 km sampling 1077 (Figure 6b). Finally, greater T_e is observed in the 1078 flat slab area of Chile with 400 \times 400 km sam- 1079 pling, but does not persist in larger windows. High 1080 S velocity of the lithospheric mantle above the 1081 Chilean flat slab has also been observed with local 1082 tomography [Wagner et al., 2006], suggesting that 1083 corresponding high rigidity in Figure 6a is real. In 1084 subsequent work we discuss two possible reasons 1085 for the relationship between high S velocities and 1086 T_e at flat subduction zones (Pérez-Gussinyé et al., 1087 manuscript in preparation, 2007). 1088

[56] Figure 8 summarizes the relationship between 1089 T_e and S velocities. Here we show a scatterplot of 1090 T_e estimates versus shear wave velocity at 100 km 1091 depth (red dots) sampled from the maps shown in 1092 Figure 6. Superimposed, we also show the mean 1093 value of T_e within 10 km bins versus the cor- 1094 responding mean of S-velocity at 100 km (blue 1095 squares). Although the standard deviation of 1096 S-velocities in the relationship is large, T_e is 1097 positively correlated with S-velocity, as already 1098 suggested from visual comparison.

[57] There are numerous possible contributors to 1100 the scatter in the relationship. First and foremost, T_e 1101 and shear wave velocity are measuring two very 1102 different things. T_e is fundamentally an integral of 1103 the bending stress that supports lithospheric loads, 1104 and as such it is most sensitive to parameters of the 1105 power law creep rheology (i.e., temperature, min- 1106 eral composition, water, and to a lesser degree 1107 strain rate) integrated over the entire crust and the 1108 uppermost mantle [e.g., Lowry and Smith, 1995]. 1109 Shear wave velocity is primarily a measure of rock 1110 shear rigidity (with a weak dependence on density) 1111 that is most sensitive to temperature, mineral 1112 composition, water and the presence of partial 1113 melt. The dependencies on temperature, mineral 1114 composition and water content have similar sign 1115 for both T_e and S velocity, which is responsible for 1116 the positive correlation between the two. However, 1117 the relative sensitivities to each of these fields are 1118





Figure 8. (a) Red dots are scatterplots of T_e estimates obtained using 400 × 400 km windows (Figure 6a) versus shear wave velocity at 100 km depth from Feng et al. (submitted manuscript, 2007) (Figure 6f). Blue squares represent the mean of T_e within 10 km bins versus mean S-velocities; bars are one standard deviation. (b and c) The same as in Figure 8a, but for T_e estimates obtained using 600 × 600 km and 800 × 800 km windows, respectively.

1119 quite different. Considering these differences and 1120 the differences in depth-sampling (a complicated 1121 integral of properties in the crust and upper mantle 1122 versus a slice at 100 km depth) and in resolution of 1123 the measurements, the scatter in the relationship is 1124 not at all surprising.

[58] Effects of differences in spatial resolution are 1125 apparent in comparisons between T_e and S-veloc- 1126 ities for different T_e estimation window sizes. T_e 1127 analyses for the 400 \times 400 km windows yield far 1128 more estimates of less than 20 km than do the 1129 larger windows (compare, e.g., Figure 8a with 1130 Figures 8b and 8c). The smaller T_e estimates 1131 derived using larger windows are found predomi- 1132 nantly in the Cordillera where both T_e and S 1133 velocity are most likely to be significantly per- 1134 turbed by non-thermal effects. Simultaneously, 1135 smaller T_e that may be present in weak lineaments 1136 separating the cratons tends to be averaged out, 1137 resulting in a weaker correlation of T_e and S 1138 velocity for low T_e and larger windows. On the 1139 other hand, at the high end of the T_e range, the 1140 standard deviation of S-velocities is larger for 1141 estimates obtained with small windows than with 1142 larger ones (compare Figure 8a with Figure 8c). We 1143 believe that this results from the inability of small 1144 windows to recover large T_e (see Figures 4 and 5). 1145

5.3. Te and Surface Heat Flow

[59] Heat flow data reflects the thermal state of the 1148 lithosphere and as such should also relate to T_e 1149 [Lowry and Smith, 1995]. However, uncertainties 1150 in the relationship of surface heat flow to temper- 1151 ature at depth can be large, as poorly known crustal 1152 heat production and dynamical effects of subduc- 1153 tion, erosion, and hydrologic flow are all important 1154 contributors [e.g., Mareschal and Jaupart, 2004]. 1155 Also, heat flow measurements in South America 1156 are relatively sparse, and many measurements 1157 acquired by mining and oil companies are unreli-1158 able [Hamza et al., 2005]. Thus any comparison of 1159 South American T_e and heat flow requires caution. 1160

[60] We compare our T_e estimates to a recent compilation of heat flow in South America by *Hamza* 1162 *et al.* [2005]. Sampling locations are shown as 1163 black triangles in Figure 6e. *Hamza et al.* [2005] 1164 find generally lower heat flow (<60 mW/m²) in the 1165 stable part of the continent than in the Andean 1166 cordillera (>70 mW/m²). This first-order pattern 1167 of heat flow variation is in accord with our T_e 1168 estimates. 1169

[61] Smaller scale variations in heat flow are also 1170 apparent in Figure 6e, which plots the measure- 1171 ments after subtracting a regional heat flow field. 1172 The regional heat flow is a polynomial representa- 1173 tion of South American heat flow which is meant 1174 to represent the first-order westward increase from 1175 60 mW/m^2 in the Stable Platform to 70 mW/m² in 1176 the Andes [*Hamza et al.*, 2005]. Within the stable 1177

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1178 platform, higher than normal heat flow is observed 1179 near the southwestern part of the Transbrasiliano 1180 lineament and in the Amazon basin, both of which 1181 have lower T_e and shear wave velocities than the 1182 surrounding cratons (Figure 6) [*Feng et al.*, 2004; 1183 *Heintz et al.*, 2005; *Schimmel et al.*, 2003; Feng 1184 et al., submitted manuscript, 2007]. Correlation of 1185 high heat flow with low S velocity at 100 km depth 1186 suggests a correspondingly high geothermal gradi-1187 ent to lithospheric depths, resulting in the weak 1188 lithosphere observed in our T_e maps.

1189 [62] Additionally, high heat flow is observed 1190 between the northern end of the Paraná basin 1191 and the southern end of the Sao Francisco craton 1192 (Figure 6e). However, this area does not exhibit 1193 low shear wave velocity or small T_e , suggesting 1194 that this heat flow anomaly reflects very shallow 1195 processes. Alternatively, this anomaly may have 1196 too small a lateral scale to be imaged by conti-1197 nent-scale studies of T_e and S velocity.

1198 [63] Within the Andean cordillera, low heat flow is 1199 observed over the Chile, Peru and northern Colom-1200 bia flat subduction zones. These areas are associ-1201 ated with relatively high S velocity and T_e , 1202 suggesting that here also low heat flow is repre-1203 sentative of relatively low temperature at depth. In 1204 contrast, southern Chile (south of 30°S), the central 1205 Andes (between ~14° and 25°S) and Ecuador 1206 (~0°S) all have relatively high heat flow, low S 1207 velocity and low T_e (Figure 6). These are areas of 1208 normal-angle, as opposite to flat, subduction may 1209 thus have legitimately higher temperatures at shal-1210 low depth.

1212 5.4. Te and Intracontinental Seismicity

1213 [64] Intracontinental seismicity, far from plate 1214 boundaries, is thought to occur along zones such 1215 as failed rifts weakened by previous tectonic activ-1216 ity [e.g., *Sykes*, 1978], in areas where crustal 1217 inhomogeneities occur, for example large magmatic 1218 intrusions like those observed near the New Madrid 1219 seismic zone [*Campbell*, 1978], in areas of high heat 1220 flow, as also observed in the New Madrid seismic 1221 zone [*Liu and Zoback*, 1997], and where seismic 1222 velocities at depth are low [*Assumpçao et al.*, 2004].

1223 [65] In Figure 6 we compare our T_e estimates to the 1224 intracontinental seismicity based on a catalogue of 1225 shallow (\leq 45 km) earthquakes in Brazil, Paraguay 1226 and Uruguay. The catalogue presented here is an 1227 updated version of that published by *Assumpçao et* 1228 *al.* [2004] (M. Assumpçao, personal communica-1229 tion, 2006). Because of the lack of seismic stations over part of the Brazilian territory this catalogue is 1230 only valid for the region east of 65°W and south of 1231 6°N. The catalogue includes historical and instru-1232 mental data from 1861 to present and has been 1233 normalized to eliminate concentrations of events 1234 resulting from a greater population density or a 1235 larger distribution of seismic stations in particular 1236 areas (see *Assumpçao et al.* [2004] for a detailed 1237 explanation on the generation of the catalogue). 1238

[66] Assumpçao et al. [2004] found a strong cor- 1239 relation between low seismic velocities and seis- 1240 micity in southeastern Brazil, in an area located 1241 between 26° to 12°S and 58° to 43°W. Figure 6 1242 suggests that the correlation between seismic veloc- 1243 ities, seismicity and additionally low Te extends at 1244 broad-scale to the whole of Brazil, Paraguay and 1245 Uruguay. For example, seismicity concentrates west 1246 of the Paraná basin (in the area of the Asuncion dyke 1247 swarms, $\sim 27^{\circ}$ S), along the Transbrasiliano linea- 1248 ment, and along the Amazonian basin, where T_e and 1249 shear wave velocity are low. Interestingly, Te values 1250 along the Transbrasiliano lineament are low only in 1251 seismically active areas (Figure 6a). Because seis- 1252 mically active areas with low T_e are surrounded by 1253 rigid cratons, where seismicity is very sparse, we 1254 infer that the latter are strong enough to inhibit 1255 tectonism, leading to repeated focusing of deforma- 1256 tion within the younger, weaker areas. 1257

[67] The correlation between areas of low shear 1258 wave velocity, high heat flow and intracontinental 1259 seismicity probably implies that the topography of 1260 the lithosphere-asthenosphere boundary might play 1261 an important role in focusing seismicity within 1262 continental interiors, as proposed by Assumpçao 1263 et al. [2004] for southeast Brazil and Liu and 1264 Zoback [1997] for the New Madrid seismic zone 1265 in North America. These authors suggest that 1266 where the lithosphere is thinner than the surround- 1267 ing areas, the geothermal gradient is higher, result- 1268 ing in weakening of the mantle lithosphere and 1269 lower crust with respect to the surrounding regions. 1270 This leads to focusing of the far-field tectonic 1271 stress and concentration of seismicity within the 1272 brittle, upper part of the thinner lithosphere 1273 [Assumpçao et al., 2004]. 1274

[68] We note that present-day seismic activity, as 1275 evidenced by intracontinental earthquakes, tends to 1276 focus in the same areas where past tectonic activity 1277 localized, e.g., west of the Parana flood basalts, 1278 along the Transbrasiliano lineament, the Amazo- 1279 nian basin and the Tacutu graben. Thus the differ- 1280 ences in lithospheric flexural rigidity, thickness and 1281 geothermal gradients observed today between 1282



greater thickness and more depleted compositions 1335 of older terranes (>1.5 Ga old).

[72] Within the Precambrian basement of the stable 1337 platform, T_e variations are observed at regional 1338 scale: relatively lower T_e occurs at sites that have 1339 been repeatedly reactivated through geological 1340 history as major sutures, rift zones and sites of 1341 hot spot magmatism. Today, these low T_e areas 1342 concentrate most of the intracontinental seismicity 1343 and generally exhibit lower shear wave velocities 1344 and higher heat flow than the surrounding rigid 1345 cratons. It follows that cratonic interiors are strong 1346 enough to inhibit tectonism and that intracontinen- 1347 tal deformation repeatedly focuses within thin, hot 1348 and hence weak lithosphere. Finally, the recurrence 1349 of the deformation throughout the Phanerozoic in 1350 the same low T_e areas probably implies that the 1351 differences in lithospheric flexural rigidity, thick- 1352 ness and geothermal gradients observed today 1353 between these areas and the surrounding cratons 1354 are not transient, and must have been maintained, 1355 in some cases, for at least \sim 500 m.y. 1356

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1283 these areas and the surrounding cratons are not 1284 transient, and must have been maintained, in some 1285 cases, for at least ~500 m.y. These differences in 1286 lithospheric structure probably reflect a higher 1287 mantle temperature and/or volatile content early 1288 in Earth's history, which lead, during the Archean 1289 and Early/Middle Proterozoic, to the formation of 1290 thicker and more depleted lithospheres which 1291 today are inherently more resistant to deformation 1292 than lithospheres formed subsequently, as has 1293 been inferred for Europe [*Pérez-Gussinyé and* 1294 *Watts*, 2005].

1296 6. Conclusions

1297 [69] We have estimated the *Te* structure of South 1298 America by modeling the coherence of continent-1299 wide grids of topography and Bouguer gravity 1300 anomaly. The gravity anomaly was generated by 1301 combining sparsely distributed terrestrial measure-1302 ments compiled by GETECH (UK) with a 1303 uniformly distributed data set from the EIGEN-1304 CG30C model, which is a combination of terrestrial 1305 and free-air gravity derived from the satellites 1306 GRACE and CHAMP [*Foerste et al.*, 2005]. A 1307 spherical Bouguer correction was applied to the 1308 EIGEN-CG30C free-air gravity anomaly in order 1309 to preserve the true wavelength content in the 1310 Bouguer anomaly.

1311 [70] Our results demonstrate a first-order pattern 1312 in which effective elastic thickness, T_e , of the 1313 stable South American platform is much higher 1314 than that of the Andean domain. T_e correlates 1315 remarkably well with other proxies for lithospheric 1316 structure. Areas of high T_e exhibit high seismic 1317 velocities at ~100 km depth and low surface heat 1318 flow. The excellent correlation of our results with 1319 shear wave velocity maps of the continent suggests 1320 that in the future, T_e maps may not only be used to 1321 better understand lithospheric mechanical proper-1322 ties, but also to image lateral variability of litho-1323 spheric structure (e.g., thickness and composition) 1324 analogously to how seismic velocities are currently 1325 used.

¹³²⁶ [71] We also find that within the stable South Amer-¹³²⁷ ican platform, the Archean and Paleo/Middle-¹³²⁸ Proterozoic cratonic nuclei exhibit large T_e , as do ¹³²⁹ the Parnaiba and Paraná basins which are thought ¹³³⁰ to be underlain by old (~2 Ga) basement [*Cordani* ¹³³¹ *et al.*, 1984; *Brito Neves et al.*, 1984]. However, ¹³³² the Phanerozoic Patagonian terrane has generally ¹³³³ lower T_e . Thus T_e is to first order dependent on ¹³³⁴ terrane age. This dependence probably reflects the

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